# **PHYS 1P22/92** Prof. Barak Shoshany Spring 2024

22. Magnetism

# 22.1 Magnets

# North and south poles

- Interaction between magnets: like poles repel and unlike poles attract.
- Similar to the sign of electric charge, but not the same.



Likes repel

# Magnetic materials

 Magnets also attract iron and other magnetic materials which are not themselves magnets.



# Earth's magnetic field

- Compass: north pole of a magnet points towards the north **geographic** pole.
- The Earth is like a huge magnet, with the south magnetic pole at the north geographic pole.



# Magnetic dipoles

- Dipole = two poles.
- Poles cannot be separated.
   Splitting a dipole always results in two smaller dipoles.
- At the lowest scales, the individual atoms are dipoles.
- In fact, even subatomic particles (electrons, protons, neutrons) are also dipoles.



# Magnetic monopoles

- Monopole = one pole.
- Isolated north or south poles would behave like electric charges.
- Predicted by some very speculative theories, but never observed (those theories are probably incorrect).
- It's currently unknown whether magnetic monopoles exist.



# 22.2 Ferromagnets and Electromagnets

# Ferromagnetic materials

- Respond strongly to magnets.
- Can be made into magnets (temporarily or permanently).
- Examples: iron, cobalt, nickel, gadolinium, neodymium.



# Magnetization & domains



# Demagnetization

- A permanent magnet can be demagnetized by hard blows or heat.
- Ferromagnetic materials cannot be magnetized above the Curie temperature.
  - For example: 1043 K (770 °C) for iron.

# Electromagnets

- Electrical current can create temporary magnets.
- Applications:
  - MRI
  - Data storage (but not so much today)



#### Electromagnet vs. permanent magnet



# Combining electromagnet and ferromagnet



### All magnets are due to electric currents!



Atom

Electron

# Simulation

• Simulation of magnets and electromagnets:

<u>https://phet.colorado.edu/sims/cheerpj/faraday/latest/faraday.ht</u> <u>ml?simulation=magnets-and-electromagnets</u>

# 22.3 Magnetic Fields and Magnetic Field Lines

# Magnetic field

- Vector **B** at each point.
- Has the same direction as a small compass at each point.
- Field lines: connecting all the arrows (like electric field lines).





Circular current loop (similar to bar magnet)

Long straight wire (field lines circle the wire) Field goes into the screen (use right-hand rule)

Rules for magnetic field lines:

1. Tangent to the magnetic field at each point.

2. Density of lines is proportional to magnitude of the magnetic field.

3. Can never cross each other (why?)

### 4. Always form closed loops from north to south pole.

(if magnetic monopoles existed, lines would start and end on them.)

### Demonstration

See the magnetic field in 3D! (Live demonstration)

22.4 Magnetic Field Strength: Force on a Moving Charge in a Magnetic Field

#### **Cross product**:

#### a × b

- Direction given by right-hand rule.
- Magnitude given by area of a parallelogram:

 $|\mathbf{a} \times \mathbf{b}| = |\mathbf{a}||\mathbf{b}| \sin \theta$ 

 $\theta$  is the angle between **a** and **b**.

Properties (prove!):

- $\mathbf{a} \times \mathbf{a} = \mathbf{0}$
- $\mathbf{a} \times \mathbf{b} = \mathbf{0}$  if  $\mathbf{a}$  parallel to  $\mathbf{b}$
- $|\mathbf{a} \times \mathbf{b}| = |\mathbf{a}||\mathbf{b}|$  if **a** perpendicular to **b**
- $\mathbf{a} \times \mathbf{b} = -\mathbf{b} \times \mathbf{a}$



# Force on a point charge (electric + magnetic)

- Lorentz force:  $\mathbf{F} = q\mathbf{E} + q\mathbf{v} \times \mathbf{B}$
- $\mathbf{F} = \text{force}$
- q = charge
- $\mathbf{E} = \text{electric field}$
- **v** = velocity of charge
- **B** = magnetic field
- × = cross product



# Force on a point charge (magnetic only)

• Lorentz force:

 $\mathbf{F} = q\mathbf{v} \times \mathbf{B}$ 

- Direction: right-hand rule (there are two, either one works)
- Magnitude:

 $F = qvB\sin\theta$ 

- Properties (prove!):
  - No force on neutral particle.
  - No force on charge at rest.
  - No force on charge moving parallel to magnetic field.
  - Strongest force on charge moving perpendicular to magnetic field.



$$F = qvB\sin\theta \implies B = \frac{F}{qv\sin\theta}$$

• SI units: tesla (T)

$$1 T = \frac{1 N}{C \cdot m/s} = \frac{kg}{A \cdot s^2} \text{ (in base SI)}$$

• Examples:

- Earth's magnetic field:  $\sim 32 \ \mu T = 32 \times 10^{-6} \ T$
- Refrigerator magnet:  $\sim 5 \text{ mT} = 5 \times 10^{-3} \text{ T}$
- MRI machine:  $\sim 1.5$  to 3 T
- Large Hadron Collider: ~8 T
- White dwarf, a very dense stellar remnant (dead star):  $\sim 100 \text{ T}$
- Neutron star, an even denser stellar remnant:  $\sim 10^4$  to  $10^{11}$  T (magnetar)
- Old units: gauss (G),  $1 \text{ G} = 10^{-4} \text{ T}$ .

22.5 Force on a Moving Charge in a Magnetic Field: Examples and Applications

# **Circular** motion

- Magnetic force is always perpendicular to velocity (cross product).
- So it does **no work** on the charged particle:

$$\theta = \frac{\pi}{2} \Rightarrow$$

$$W = \mathbf{F} \cdot \mathbf{d} = Fd \cos \theta = 0$$

• Direction of motion changes, but not speed (kinetic energy is constant).



### **Circular** motion

• If **v** is perpendicular to **B** then  $F = qvB \sin \theta = qvB$ 

 $(\theta = \pi/2, \text{ not the same } \theta!)$ 

• This supplies the centripetal force:

$$F = \frac{mv^2}{r} = qvB$$

• Thus:

$$r = \frac{mv}{qB}$$



# Spiral motion

- If v is not perpendicular to B then we consider its component  $v_{\perp}$  perpendicular to the field.
- The component  $v_{\parallel}$  parallel to the field is unaffected.
- This causes spiral motion instead of circular.
- Application: bubble chamber. Can detect particles and measure their mass and charge by observing their path in a magnetic field.



# Magnetic mirror

- Field strength increases in direction of motion.
- Charges slow and eventually reverse their motion.



# Auroras (northern/southern lights)

- Solar winds are very fast and very hot charged particles that emanate from the Sun.
  - Protons
  - Electrons
  - Alpha particles (2 protons + 2 neutrons, like a <sup>4</sup>He nucleus)
- The Earth is shielded from them by its magnetic field at the magnetosphere.
- Otherwise the atmosphere would be stripped off the planet!



The field is "squashed" on the side of the Sun due to the solar winds.

.........

- The particles flow along the magnetic field lines towards the poles.
- When they collide with nitrogen and oxygen atoms in the atmosphere, causing them to emit photons.
  - Oxygen = green/yellow/red light
  - Nitrogen = blue light
- This is called an aurora and can be seen at the "auroral oval", around 10-20° from the poles.





Forecast Lead Time: 64 minutes HPI: 31.6 GW (Range 5 to 200)

IOAA

NOAA Space Weather Prediction Center Aurora Forecast For 2022-03-10 03:39 (UTC)

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Forecast Lead Time: 64 minutes HPI: 30.1 GW (Range 5 to 200)

Probability of Aurora10%50%90%

0 1 2 3 4 >4 Approximate Energy Deposition ergs/cm2 OVATION Aurora Model Model Run at 2022-03-10 02:35 (UTC) L1 Observations at 2022-03-10 02:31 (UTC)

Probability of Aurora10%50%90%

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### The Van Allen radiation belts

- Regions where energetic charged particles from the solar wind are trapped.
- Inner belt: 1,000-12,000 km from Earth
- Outer belt: 13,000 to 60,000 km from Earth



# (Skipping sections 22.6-22.8)

# 22.9 Magnetic Fields Produced by Currents: Ampere's Law

# Long straight wire

$$B = \frac{\mu_0 I}{2\pi r}$$

- B = magnitude of magnetic field.
- $\mu_0 \approx 1.26 \times 10^{-6} \text{ N/A}^2 = \text{vacuum}$ permeability.
- I = current.
- r = radius (shortest distance).
- Direction determined by the righthand rule.



# Circular loop



Only valid at the center of the loop.



# Solenoid (long coil of wire)

Approximately uniform field:  $B = \mu_0 nI$ 

*n* is the number of loops per unit length. If N = number of loops and L = length, then  $n \equiv N/L$ .

